## Geek Week, Spring 2016

You think a "day" should be shorter than a "year", and only one sunrise per day, right?

# Two Sunrises on Mercury <br> -- Count 'em TWO! 

(And, a "day" longer than the "year"??)

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## Start with Earth - because we know it!

- Rotation on axis (n-axis) is 360 deg/day (pretty fast)
- If we were not moving around the Sun, we would see the Sun rise in the East and set in the West ...right?
Dr. Scott, please demonstrate ... (so n-axis = E->W)
- Suppose we did not turn on our axis - where would the sun rise and set as we orbited (n-orbit)? Dr. Scott, we're waiting!
- So, the effect of the orbit is a "west to east" motion of the Sun - interesting! (n-orbit = W-E !!)
- So, if the rate due to the orbit is bigger than the axis rate ... the sun moves "differently" in the sky (West to East)


## But, n-orbit not quite that simple ...

- All the planets have elliptical orbits
- Mercury is more eccentric than other inner planets
- At perihelion, the planet moving faster, so the n-orbit is higher
- At aphelion, planet slower, thus n-orbit smaller
- Let's call them n-per and n-ap (both of these are basically " $n$-orbits")
- Remember, if the rate due to the orbit (now n-per, nap ) is bigger than the axis rate ... the sun moves "differently" in the sky (West to East)
- Thus .. If the n-axis rate is *between* those two ....


## Data Chart for all the planets

| Planet | ecc | $\begin{gathered} \mathbf{a} \\ (A U) \\ \hline \end{gathered}$ | Orbit Period (days) | mean (deg/day) | Axis period (days) | $\begin{gathered} \mathrm{n} \text {-axis } \\ \text { (deg/day) } \end{gathered}$ | n-per | n-ap | Nper>Naxis>Napp? |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Mercury | 0.206 | 0.39 | 87.97 | 4.09 | 58.81 | 6.12 | 6.35 | 2.76 | YES!!! |
| Venus | 0.007 | 0.72 | 224.70 | 1.60 | -243.69 | -1.48 | 1.62 | 1.58 | no - backwards! |
| Earth | 0.017 | 1.00 | 365.25 | 0.99 | 1.00 | 360.00 | 1.02 | 0.95 | no |
| Mars | 0.093 | 1.52 | 868.98 | 0.41 | 1.03 | 350.89 | 0.50 | 0.35 | no |
| Jupiter | 0.048 | 5.20 | 4331.87 | 0.08 | 0.41 | 870.53 | 0.09 | 0.08 | no |
| Saturn | 0.056 | 9.54 | 10760.27 | 0.03 | 0.44 | 822.86 | 0.04 | 0.03 | no |
| Uranus | 0.047 | 19.18 | 30684.65 | 0.01 | 0.72 | 501.16 | 0.01 | 0.01 | no |
| Neptune | 0.009 | 30.06 | 60189.55 | 0.01 | 0.67 | 536.31 | 0.01 | 0.01 | no |
| Pluto | 0.248 | 39.53 | 90465.12 | 0.00 | 6.41 | 56.21 | 0.01 | 0.00 | no |

ecc=eccentricity - how circular (0) or elliptical ( $>0$ ) is the orbit?
a = semi-major axis - rough distance from sun
orbit period = time for planet to orbit once around the Sun
mean $=$ mean daily motion $=360$ degrees $/$ orbit period
But, depending on eccentricity, can be higher or lower at different points axis period = fixed rotation on axis (how fast is the planet turning on axis) n-axis $=$ daily rotation rate on axis $($ fixed $)=360 /($ axis period $)$

Mercury has more eccentric orbit .. N-orbit changes at perihelion/aphelion!!
n-per = daily motion at perihelion (close approach distance to Sun) > n-ap
n-ap = daily motion at aphelion (far approach distance to Sun) <n-per

## Earth and Beyond ... if n-axis < n-perihelion ... Doh!



## Mercury/Venus ... if n-axis < n-perihelion ...



## Mercury ... what does this mean?

- Mercury orbits the sun faster at the perihelion (close approach) compared to the aphelion
- The normal axis rotation (n-axis gives the East to West motion of the sun) is a little slower than the orbit rate at perihelion (n-per gives the West to East motion)
- This means as Mercury approaches the perihelion location - the Sun will momentary "stop" in the sky, then reverse a little, then go forward again!
- If this happens to occur around the sunrise time ... double-sunrise, baby!
- Let's plot angle above the horizon (for center of Sun) against days near the perihelion ...


## Ta-Da ??



## Oh - there we go!



## How much of the sun is involved?

- Graph plots the center of the Sun (so for the leading edge - it would actually occur a little earlier)
- At Mercury's distance - the Sun has an angular size of about 1.73 degrees across the diameter - thus center to edge is about 0.86 degrees
- The center rises a maximum of 0.55 degrees $+0.86=$ $\sim 1.4$ degrees, so not all Sun above horizon (more than half shows then sinks again, then rises)


## Let's put some Suns on the graph!



Green line for the grassy (?) horizon!

## Will Mercury always do this?

- Yes and no ... there is a critical eccentricity ...
- Mercury's eccentricity varies between about 0.12 and 0.23 (currently 0.206)
- If the eccentricity is exactly 0.191059 .. Sun would just momentarily "stop" and then go forward (but no doublesunrise, just a paused one!)
- If eccentricity below 0.191059 - n-axis > n-per and back to a boring single sunrise per day
- Currently Mercury's eccentricity on the rise - so many more years of double-sunrises!


## Acknowledgments

- Calculations from Astronomy Morsels III by Jean Meeus
- If you want to do astronomy calculations - he is THE source of information!
- Graphs courtesy of Microsoft Excel
- Presentation by ... um, oh yeah, Powerpoint
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- Thanks to the Arts and Sciences Seminar committee for letting me do my Astro-gig once a year!


## Mercury - "Day" longer than "Year"?

- Mercury orbit the sun $=87.98$ days $\sim 90$ days
- Mercury period on axis $=58.6$ days $\sim 60$ days
- Every 15 days, 90 degrees on axis (look for colors indicating full rotation on axis)
$-1 / 6^{\text {th }}$ way around $\operatorname{sun}$ (60 degrees, $120,180,240,300,360$, etc)
Year 1!
Year 2!!
Day 0-0 deg - noon (W)

Day 15 - ccw 60 (S)
Day 30 - ccw 120 (E)
Day 45 - ccw 180 (N)
Day 60 - ccw 240 (W)
Day 75 - ccw 300 (S)
Day 90 - ccw 360 (E)

Day 105 - ccw 60 (N)
Day 120 - ccw 120 (W)
Day 135 - ccw 180 (S)
Day 150 - ccw 240 (E)
Day 165 - ccw 300 (N)
Day 180 - ccw 360 (W) 14

